

Gravitational Wave Discovery

The background of the slide features a visualization of gravitational waves on the left, showing concentric ripples in shades of orange, yellow, and red. On the right, there is a realistic image of the Earth as seen from space, showing continents and oceans. The overall background is dark, suggesting a cosmic or space-themed setting.

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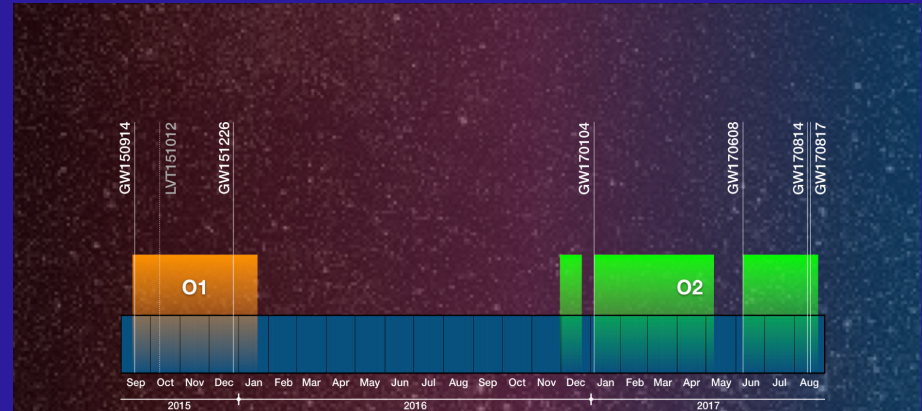
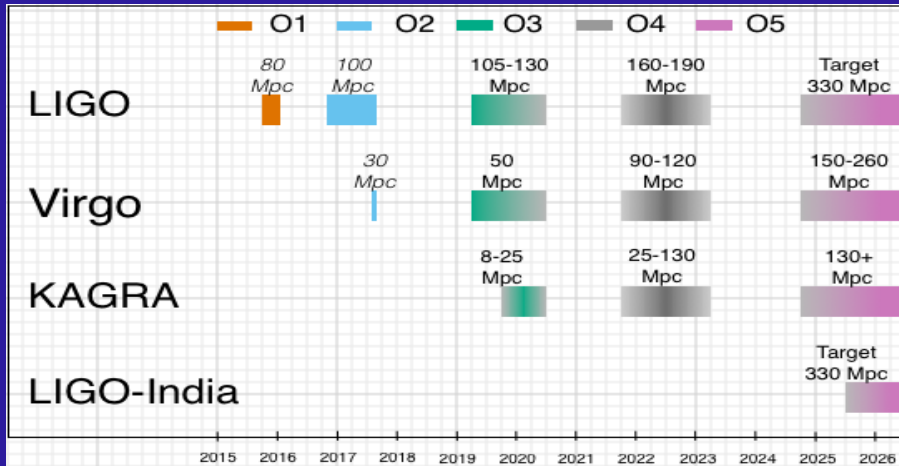
**Gravitational Wave Astronomy Group
Department of Physics, UWA**

ARC Center of Excellence for Gravitational Wave Discovery (OzGrav)

**Gravitational Wave Technology and Education Research Cluster
Machine Learning Application for Physical Sciences Research Cluster**

LIGO Scientific Collaboration (LSC)

Exciting Time for GW Discovery



O1/O2 Discoveries:

- 11 confirmed detections: 10 BBHs (4 previously unpublished), 1 BNS
- 14 marginal detections ($FAR < 1/30$ days, $P_{astro} < 50\%$)

GraceDB — Gravitational-Wave Candidate Event Database

HOME PUBLIC ALERTS SEARCH LATEST DOCUMENTATION LOGIN

LIGO/Virgo Public Alerts

Detection candidates: 22 (O3: <https://gracedb.ligo.org>)

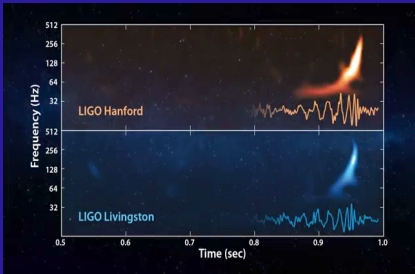
SORT: EVENT ID (A-Z)

Event ID	Possible Source (Probability)	UTC	GCN	Location	FAR
S190808ae	Terrestrial (57%), BNS (43%)	Aug. 8, 2019 22:21:21 UTC	GCN Circulars Notices VQE		1.0622 per year
S190728a	BBH (95%), MassGap (5%)	July 28, 2019 06:45:10 UTC	GCN Circulars Notices VQE		1 per 1.2541e+15 years
S190727h	BBH (92%), Terrestrial (5%), MassGap (3%)	July 27, 2019 06:03:33 UTC	GCN Circulars Notices VQE		1 per 229.92 years
S190720a	BBH (99%), Terrestrial (1%)	July 20, 2019 00:08:36 UTC	GCN Circulars Notices VQE		1 per 8.3367 years
S190718y	Terrestrial (98%), BNS (2%)	July 18, 2019 14:35:12 UTC	GCN Circulars Notices VQE		1.1514 per year
S190707a	BBH (>99%)	July 7, 2019 09:33:26 UTC	GCN Circulars Notices VQE		1 per 6018.9 years

O3 Discoveries:

NSBH: 5
BNS: 3
Mass Gap: 2
BBH: 21

• 2015: First Detection of GWs from Binary Black Hole Merger



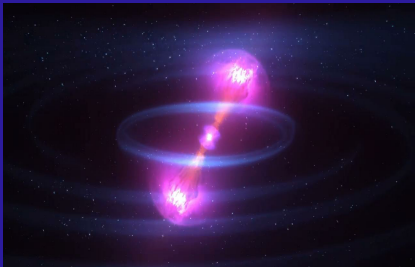
• 2017: Nobel Prize in Physics



Rainer Weiss (MIT) Kip Thorne (Caltech) Barry Barish (Caltech)

• 2017: First Detection of GWs and Light from Binary Neutron Star Merger

• 2019 Apr 1 – : First Open Public Alert



UWA Gravitational Wave Astronomy Group At the Frontier of GW Discoveries



User Guide

Primer on public alerts for astronomers from the LIGO and Virgo gravitational-wave observatories.

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Online Pipelines

A number of search pipelines run in a low latency, online mode. These can be divided into two groups, modeled and unmodeled. The modeled ([CBC](#)) searches specifically look for signals from compact binary mergers of neutron stars and black holes ([BNS](#), [NSBH](#), and [BBH](#) systems). The unmodeled (Burst) searches on the other hand, are capable of detecting signals from a wide variety of astrophysical sources in addition to compact binary mergers: core-collapse of massive stars, magnetar star-quakes, and more speculative sources such as intersecting cosmic strings or as-yet unknown GW sources.

Modeled Search

[GstLAL](#), [MBTAOnline](#), [PyCBC Live](#) and [SPIIR](#) are matched-filtering based analysis pipelines that rapidly identify compact binary merger events, with $\lesssim 1$ minute latencies. They use discrete banks of waveform templates to cover the target parameter space of compact binaries, with all pipelines covering the mass ranges corresponding to [BNS](#), [NSBH](#), and [BBH](#) systems.

A coincident analysis is performed by all pipelines, where candidate events are extracted separately from each detector via matched-filtering and later combined across detectors. [SPIIR](#) extracts candidates from each detector via matched-filtering and looks for coherent responses from the other detectors to provide source localization. Of the four pipelines, [GstLAL](#) and [MBTAOnline](#) use several banks of matched filters to cover the detectors bandwidth, i.e., the templates are split across multiple frequency bands. All pipelines also implement different kinds of signal-based vetoes to reject instrumental transients that cause large [SNR](#) values but can otherwise be easily distinguished from compact binary coalescence signals.



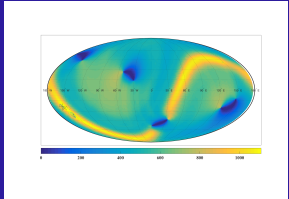
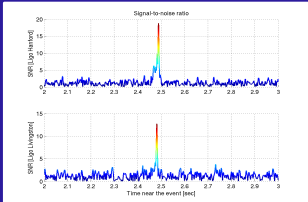
- Authorized by LVC to generate GW open public alerts
 - One of the 5 groups in the world !
 - Based on 2 UWA PhD thesis + 1 MS thesis + several research papers
- Our detections are among the fastest!
 - Send alerts within 18-30 s after binary merger
 - Now aiming at pre-merger detections
 - To facilitate prompt EM follow up observations
- Finalist for 2019 HPC Wire Reader's Choice Award for "best use of HPC in physical sciences"



UWA GWA Group: MS/PhD Projects

• Rapid Online Detection and Follow-ups of Gravitational Waves

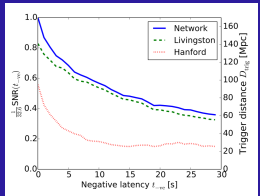
(Projects within the LIGO-Virgo Scientific Collaboration)



(GCN Notices at gracedb.ligo.org, Chu, Q. 2017 UWA PhD thesis, Hooper, S. 2014 UWA PhD thesis...)

• Pre-merger Detection of Gravitational Waves and Electromagnetic Follow-ups

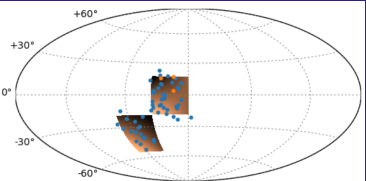
(Projects within the LIGO-Virgo Scientific Collaboration, collaborate with ICRAR-Curtin, MWA and ASKAP)



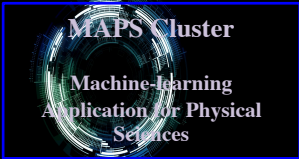
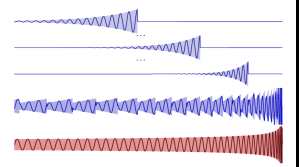
(Clancy, J. et al 2019, MNRAS Letter)

• Machine Learning for GW Discoveries

(Projects within the LIGO-Virgo Scientific Collaboration, collaborate with ICRAR-UWA and UWA Computer Science)



(Chatterjee, C. et al 2019 PRD, accepted upon revision)



UWA GWA Group: MS/PhD Projects

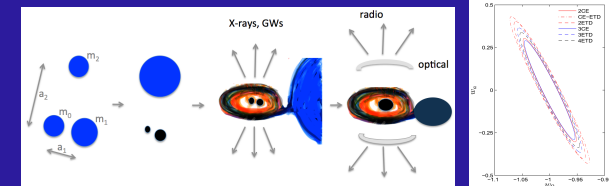
- **Coincidence Search for Gravitational Wave and Fast Radio Bursts/Gamma Ray Bursts**

(Project within the LIGO-Virgo Scientific Collaboration, collaborate with ASKAP/MWA/CRAFT)



- **Binary Black Hole Merger Modeling, and Using GW Data to Probe our Universe**

(Collaborate with Caltech and USTC)

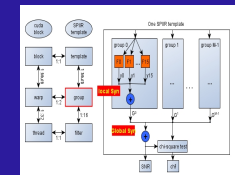
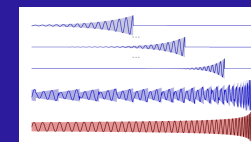


(Wen & Phinney 2018)

(Zhao & Wen 2018 PRD)

- **High-performance Computing, Algorithm Design, Mathematical Optimization, and GPU-Acceleration**

(Collaborate with UWA Computer Science and Tsinghua U in China)



(Hooper S. et al 2012, Luan et al 2012, Liu et al 2012, Guo et al 2018, Chu 2017)

